

## SIMPLE FORMULAE FOR ESTIMATING STARK BROADENING PARAMETERS OF NEUTRAL ATOM LINES

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RESUME.- Nous avons présenté formules pour estimation de la largeur et du déplacement Stark des raies spectrales des neutres.

SUMMARY.- Simple formulae for estimating Stark width and shift of neutral atom lines are presented.

In 1978, Freudenstein and Cooper [1] suggested a simple method for evaluation of electron-impact widths of neutral atom lines, based on the simplification of GBKO [2] method. We develop here this approach further and extend its applicability to the shift calculations also.

The half-half width ( $w$ ) and shift ( $d$ ) of a neutral atom spectral line broadened by electron impacts are given by [2]

$$w + id = \frac{4\pi}{3} N_e \left(\frac{k}{m\nu}\right)^2 \int \frac{d\nu}{\nu} f(\nu) \left\{ \frac{1}{2} \mathfrak{S}_{\min}^2 + \sum_i R_{i i'}^2 \left[ a_{i i'}(Z_{i i'}^{\min}) - i \varepsilon_{i i'} b_{i i'}(Z_{i i'}^{\min}) \right] + \sum_{f'} R_{f' f'}^2 \left[ a_{f' f'}(Z_{f' f'}^{\min}) + i \varepsilon_{f' f'} b_{f' f'}(Z_{f' f'}^{\min}) \right] \right\} \quad (1)$$

where  $R_{i i'}^2$  is the square of the coordinate operator matrix element,  $i, i'$  denote the initial and final states, and  $i', f'$  are the corresponding perturbing states within the dipole approximation.  $\varepsilon_{i i'} = (E_{i'} - E_i) / |E_{i'} - E_i|$ , where  $E_j$  and  $E_{j'}$  are the energies of the corresponding states. The minimum impact parameter  $\mathfrak{S}_{\min}$  allowed by the unitarity condition [2] is given by

$$\mathfrak{S}_{\min}^2 = \frac{2}{3} \left(\frac{k}{m\nu}\right)^2 \left| \sum_i R_{i i'}^2 \left[ A_{i i'}(Z_{i i'}) - i \varepsilon_{i i'} B_{i i'}(Z_{i i'}) \right] + \sum_{f'} R_{f' f'}^2 \left[ A_{f' f'}(Z_{f' f'}) + i \varepsilon_{f' f'} B_{f' f'}(Z_{f' f'}) \right] \right| \quad (2)$$

where  $a_{j j'}$ ,  $b_{j j'}$ ,  $A_{j j'}$  and  $B_{j j'}$  are the GBKO [2] Stark broadening functions of the arguments  $z_{j j'}$  and  $\tilde{z}_{j j'}$  [2], and  $\mathfrak{S}$  is the impact parameter.

In order to simplify Eqs. (1) and (2), we introduce here the approximation

$$\left| \sum_j R_{j j'}^2 \left[ A_{j j'} + i \varepsilon_{j j'} B_{j j'} \right] \right| \approx \sum_j R_{j j'}^2 \left| \left[ A_{j j'} + i \varepsilon_{j j'} B_{j j'} \right] \right| \quad (3)$$

For a series of complex numbers  $z_j$  we have  $|\sum z_j| \leq \sum |z_j|$ , where the sign of equality holds in the case when all  $z_j$  have equal arguments. This means that  $A_{j j'} \gg B_{j j'}$  which is satisfied for close collisions, high velocities or close perturbing levels, giving usually the principal contribution to the line broadening

Define  $\eta_{jj} \equiv |E_j - E_{j'}| / 3kT$ . Then

$$w + id \approx 1.089N_e \left(\frac{h\nu}{m}\right) \left(\frac{E_n}{kT}\right)^{1/2} \left\{ \sum_{ii} R_{ii}^2 [f_w(\eta_{ii}, \vec{R}_{ii}) - i \epsilon_{ii} f_d(\eta_{ii}, \vec{R}_{ii})] + \sum_{ff} R_{ff}^2 [f_w(\eta_{ff}, \vec{R}_{ff}) + i \epsilon_{ff} f_d(\eta_{ff}, \vec{R}_{ff})] \right\} \quad (4)$$

Approximate fitted expressions for  $f_w$  and  $f_d$  are:

$$f_w(x) = e^{-1.33x} \ln\left(1 + \frac{2.27}{x}\right) + \frac{0.487x}{0.153+x} + \frac{x}{7.93+x^3}$$

$$f_d(x) = 1.571 e^{-2.482x} + \frac{1.295x}{0.415+x} + \frac{0.713x}{8.139+x^3} \quad (5)$$

where  $x = \eta_{jj} \vec{R}_{jj}$ .

In Table 1, our results obtained using Eqs. (4) and (5) are compared with results from Ref. 3 and according to Ref. 1. We can see that in the simple case of  $2p^1P-5s^1S$  transition, all calculations are in agreement. In the case of He I  $2p^1P-3d^1D$  line, where we have not a dominant perturbing level, our calculations for the width agree better with BG results.

We believe that the simple formulae presented here, will be useful when astrophysicists or physicists require a large number of neutral atom line widths and shifts influenced by the Stark effect.

Table 1. - Half-half widths and shifts in Å: DK-present results; BG-results from Ref. 5 (also given in Ref. 1); FC-results calculated according to Ref. 2.  $N_e = 10^{16} \text{cm}^{-3}$

T(K)	$w_{DK}$	$w_{BG}$	$w_{FC}$	$d_{DK}$	$d_{BG}$
He I $2p^1P-5s^1S$ 4438 Å line					
5000	1.46	1.41	1.34	1.73	1.51
10000	1.77	1.57	1.64	1.49	1.43
20000	1.85	1.65	1.72	1.12	1.24
30000	1.79	-	-	0.924	-
40000	1.72	1.62	1.60	0.803	0.996
He I $2p^1P-3d^1D$ 6678 Å line					
5000	0.468	0.423	0.948	0.236	0.275
10000	0.432	0.386	0.825	0.181	0.233
20000	0.428	0.349	0.696	0.135	0.196
30000	0.389	-	-	0.111	-
40000	0.364	0.318	0.573	0.096	0.161

#### REFERENCES

- [1] Freudenstein, S.A., Cooper, J., *Astrophys. J.*, **224**(1978)1079.
- [2] Griem, H.R., Baranger, M., Kolb, A.C., Oertel, G., *Phys. Rev.* **125**(1962)177.
- [3] Bennett, S.M., Griem, H.R., University of Maryland Tech. Rept. No 71-097, 1971.